

Star Radon Transform

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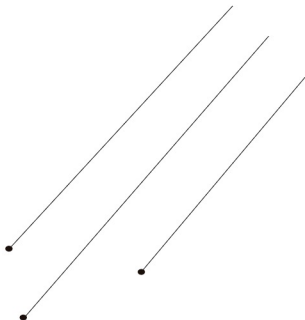
University of Arizona

Cormack Conference, Summer 2019
Joint work with Gaik Ambartsoumian.

- 1 Zhao F, Schotland J C and Markel V A 2014 Inversion of the star transform Inverse Problems 30 105001
- 2 The V-line transform with some generalizations and cone differentiation, Gaik Ambartsoumian, MJ Latifi, Inverse Problems, Vol.35 (3),2019

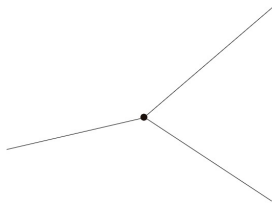
Radon Type Transforms

- **Radon Transform $\mathcal{R}$** : Integrate over lines
- **Ray Transform $\mathcal{X}_\gamma$** : Integrate over half-lines with fixed direction



Star Radon Transform

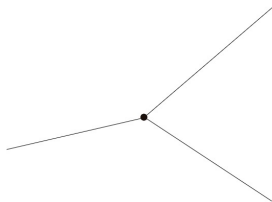
Star Radon Transform $\mathcal{S}$: Integrate over stars, i.e. sum of integrals of the function along a collection of rays emanating from the given input point.



Weighted Star Radon Transform : Weighted sum of Ray transforms

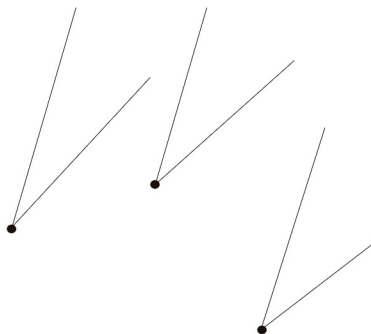
Star Radon Transform

Star Radon Transform $\mathcal{S}$: Integrate over stars, i.e. sum of integrals of the function along a collection of rays emanating from the given input point.



Weighted Star Radon Transform : Weighted sum of Ray transforms

V-line Radon Transform : When we have only two branches.



Star Radon Transform first proposed by Zhao, Schotland and Markel in [1]. Applied in imaging setups involving scattering.

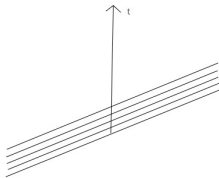
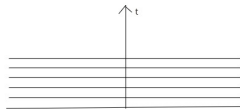
The Inversion Technique

Basic Ingredients :

- Moving Section Lemma ([2] Ambartsoumian and Latifi 2019)
- Geometric arguments involving integration and differentiation and standard Radon transform.

Moving Section Lemma

- $\int g(t) dt = c$

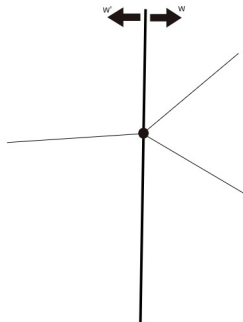


- $\int g(t) dt = \frac{1}{\sin(\theta)} c$
- Rigorous measure theoretic proof in [2]

Inversion of Star Radon Transform

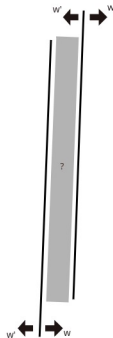
Step 1. Apply Standard Radon Transform.

When we integrate the Star Radon Transform Sf along a line, we obtain weighted integral of f on each side of the line.



Inversion of Star Radon Transform

Step 2. Find the Radon Transform of f , possibly with some weight factor depending on angle.



Theorem

Let $\mathcal{S} = \sum_{i=1}^m \mathcal{X}_{\gamma_i}$ be the star transform and let

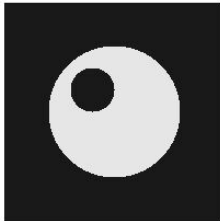
$$q(\psi) = \frac{-1}{\sum_{i=1}^m \frac{1}{\langle \psi, \gamma_i \rangle}}$$

Then the following is true for any ψ in the domain of q

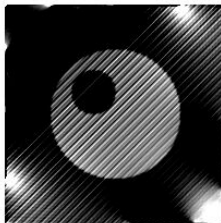
$$\mathcal{R}f(\psi, s) = q(\psi) \frac{d}{ds} \mathcal{R}(\mathcal{S}f)(\psi, s)$$

and hence if q is defined almost everywhere we can apply $\mathcal{R}^{-1}$ to get a reconstruction.

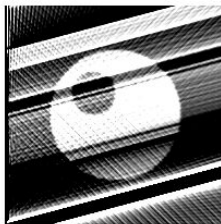
Original Phantom (300x300)



Reconstruction with branch angles : 0° , 135° , 270°



Reconstruction with branch angles : 0° , 30°



Thank you for your attention !